

# ERC in Our Research

Gia Dvali

ERC - SRNSF meeting  
(Tbilisi 2017)

In this talk I will  
mainly focus on our  
experience with ERC

# Members of ERC team:

Cesar Gomez (Professor)

Deb Sarkar } Post docs

Alexis Helou }

Kepa Sousa }

Jaba Chelidze } Ph.D.

Mischa Panchenko } Students

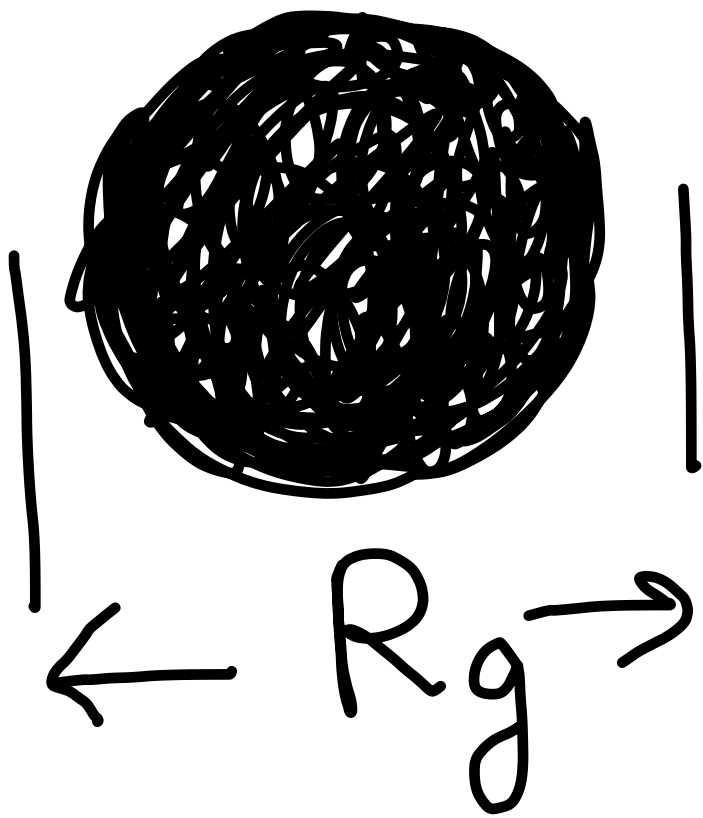
Raul Letschka }

Collaborators:

Cesar Gomez

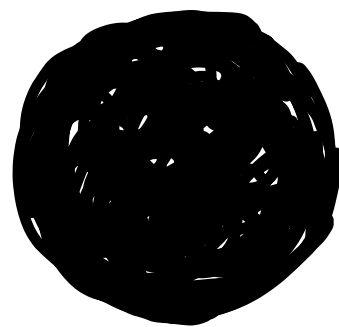
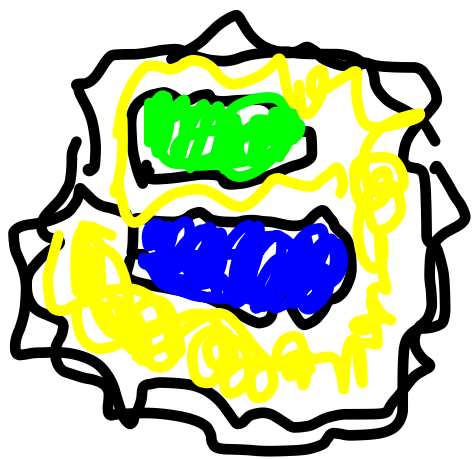
Daniel Flassig  
Alex Pritzel  
Nico Wintergerst  
Andre Franca  
Mischa Panchenko  
Dieter Lüst  
Alex Gussmann  
Artem Averin  
Sebastian Zell . . .

An object contracted  
beyond  $R_g$  becomes  
a black hole



Escape velocity = speed of light

Classically black holes  
are featureless



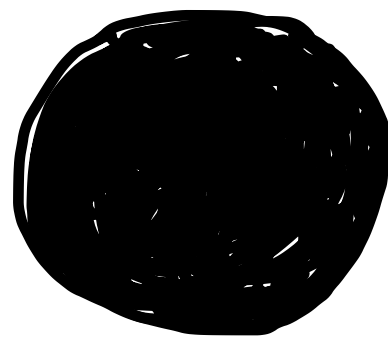
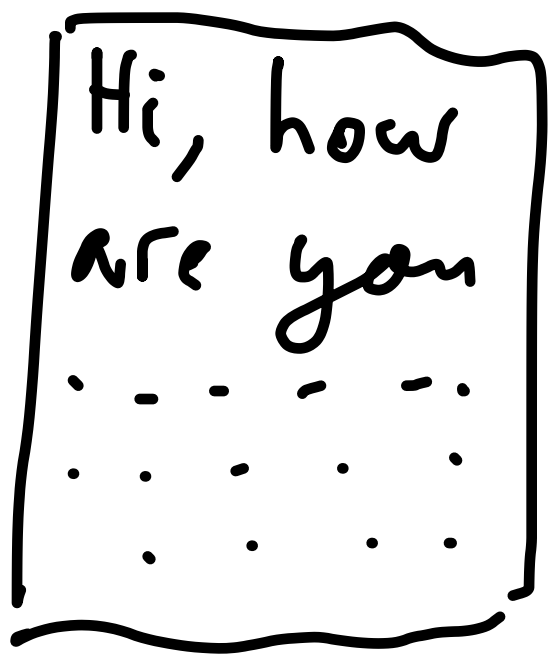
$M, J, Q$

mass

angular  
momentum

electric  
charge

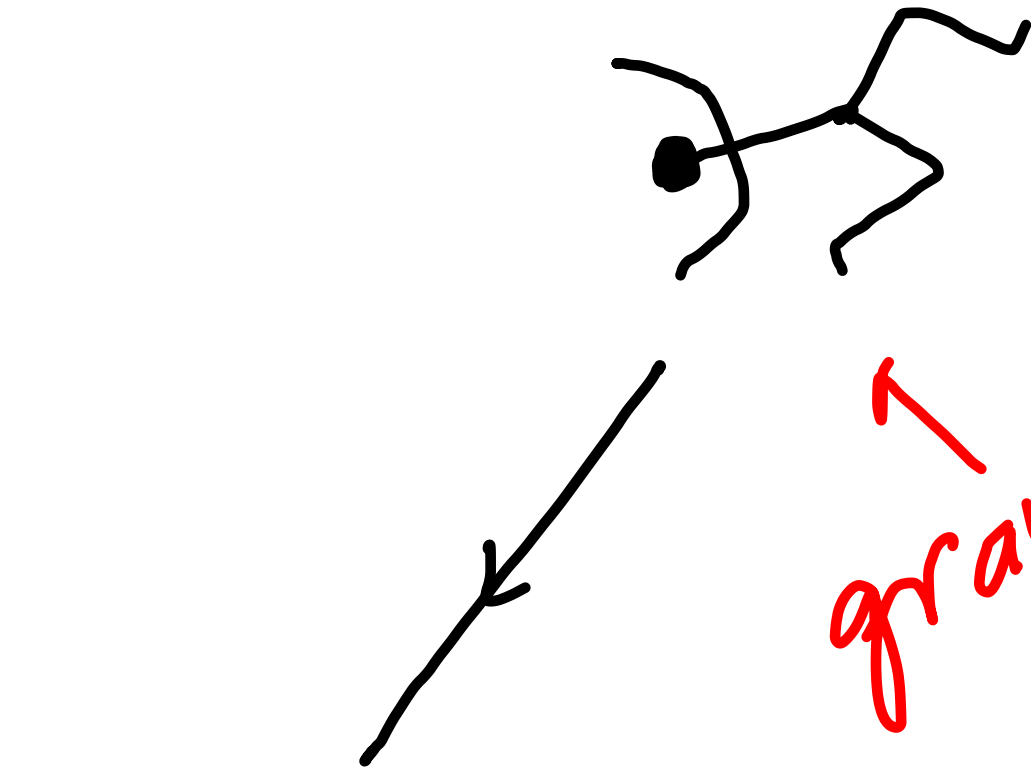
This means I cannot send  
you a message encoded  
in black hole features!



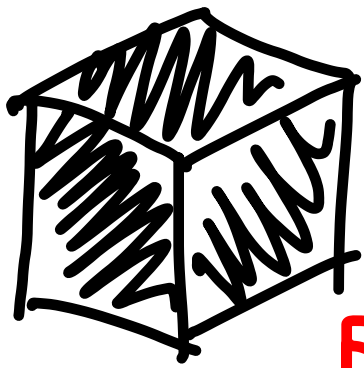
From here one would naively conclude that black holes carry very little information.

$$\ln f_0 = 0$$

But, quantum theory tells us a very different story!



↑  
gravitational  
force



Box with  
information



Consequences:

① The shortest length of nature!

Planck length

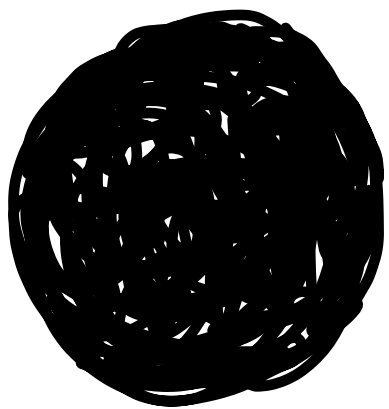
$$L_p \equiv \sqrt{\hbar G_N} \sim 10^{-33} \text{ cm}$$

# Bound on Information (Bekenstein)

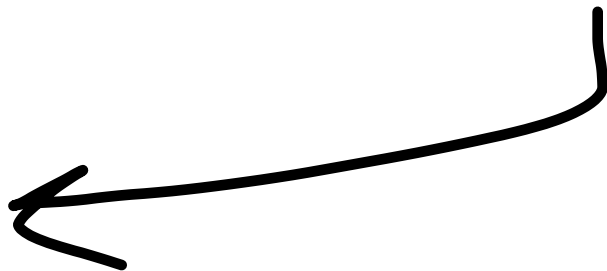
$$N_{\max} = \left( \frac{R}{L_P} \right)^2$$



← R →

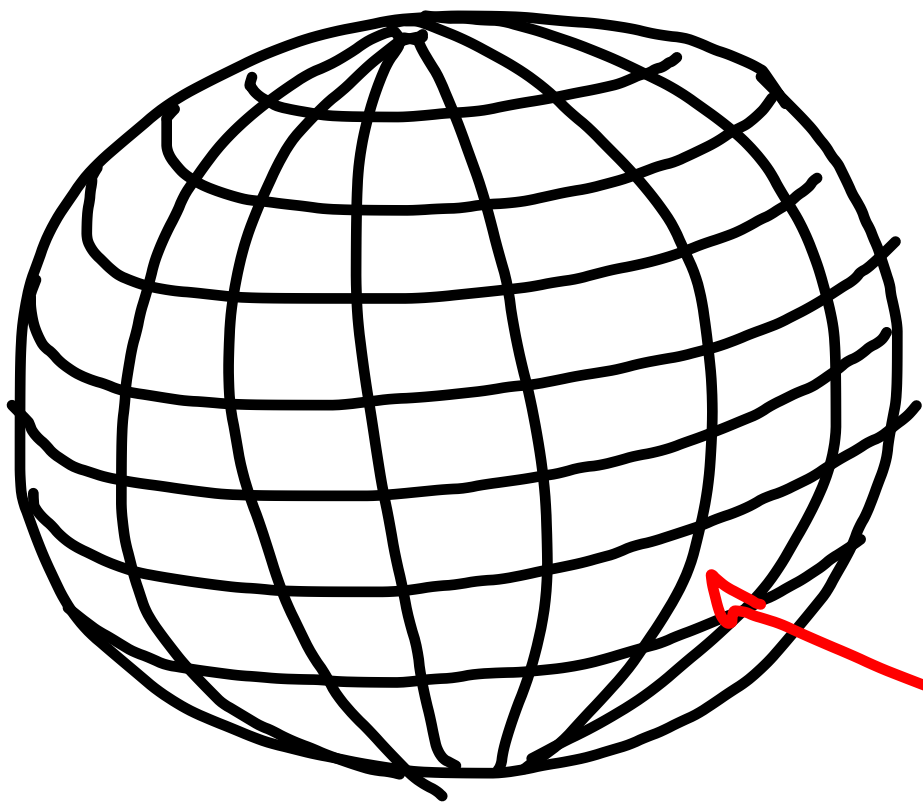


Saturated  
by black hole!



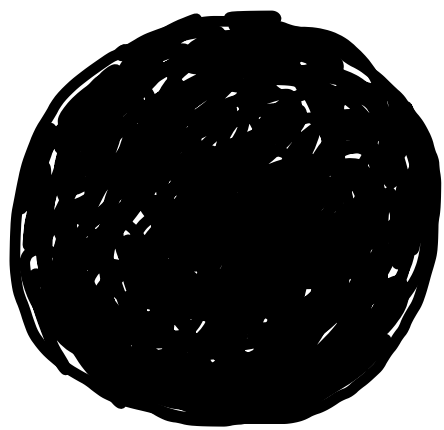
Black hole carries a maximal information (per size) given by Bekenstein entropy

$$N = \left( \frac{R_g}{L_p} \right)^2$$



Planck area pixels

But, I told you minute ago  
that classical black hole  
is featureless!

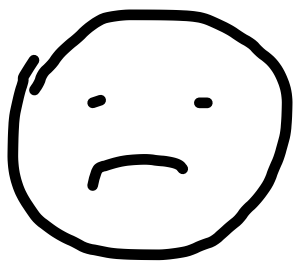


info = 0

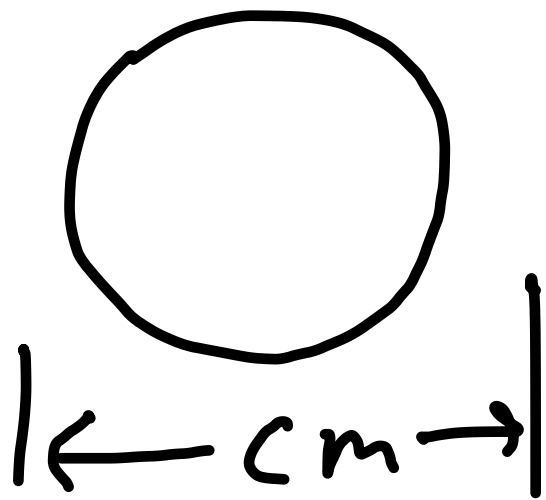


0 = ∞

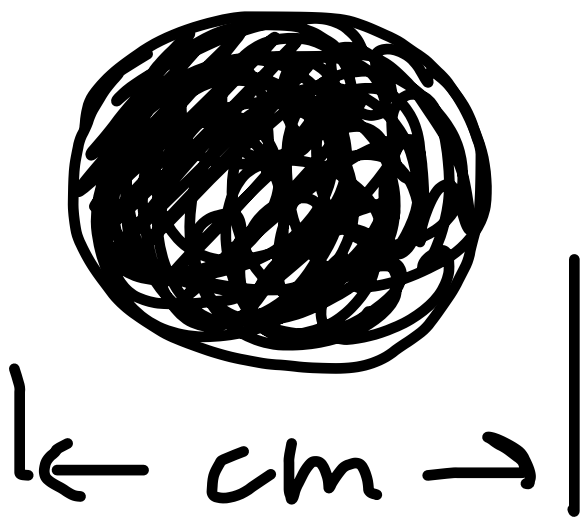
?



# Example

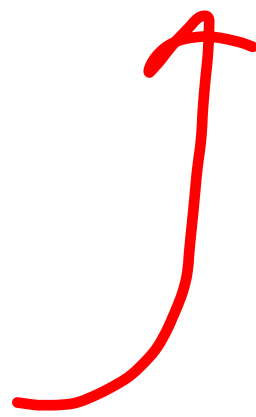


$$\Delta E \sim 10^{-4} \text{ eV}$$



$$\Delta E \sim 10^{70} \text{ eV} !$$

Black hole

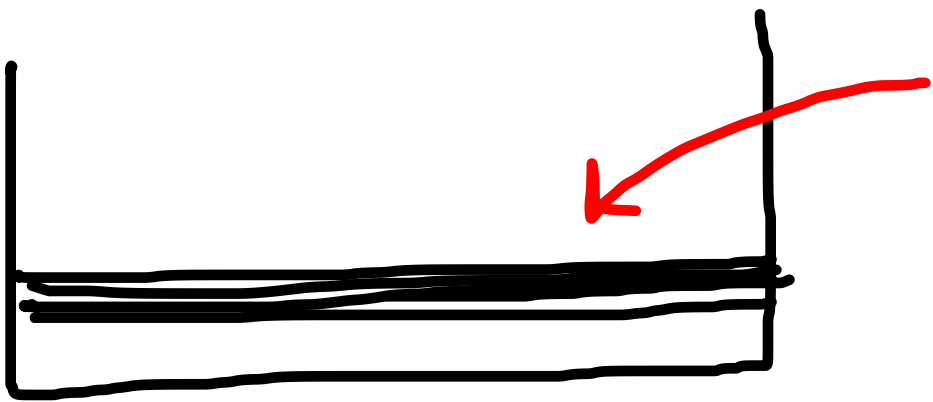


Black hole has the  
cheapest qufits!

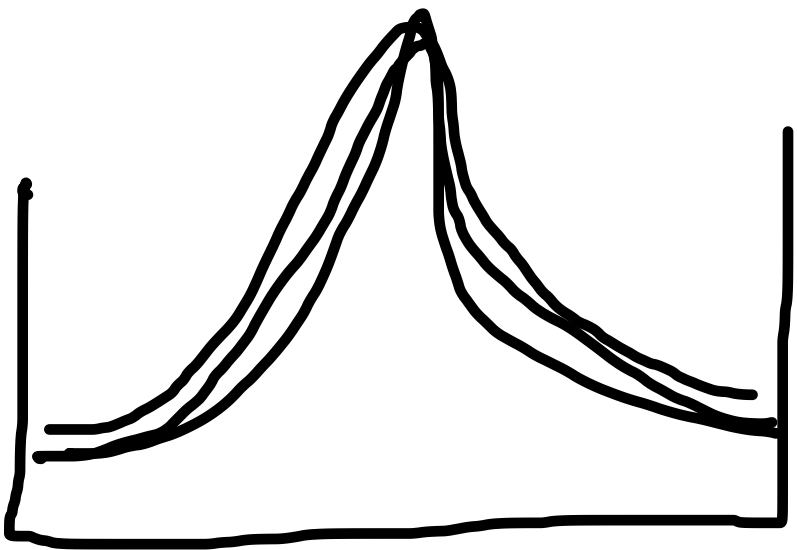
Understanding black  
holes in terms of universal  
phenomenon of quantum  
criticality of attractive  
cold bosons (gravitons)

$N$  bosons (gravitons) in  
a box

wave function



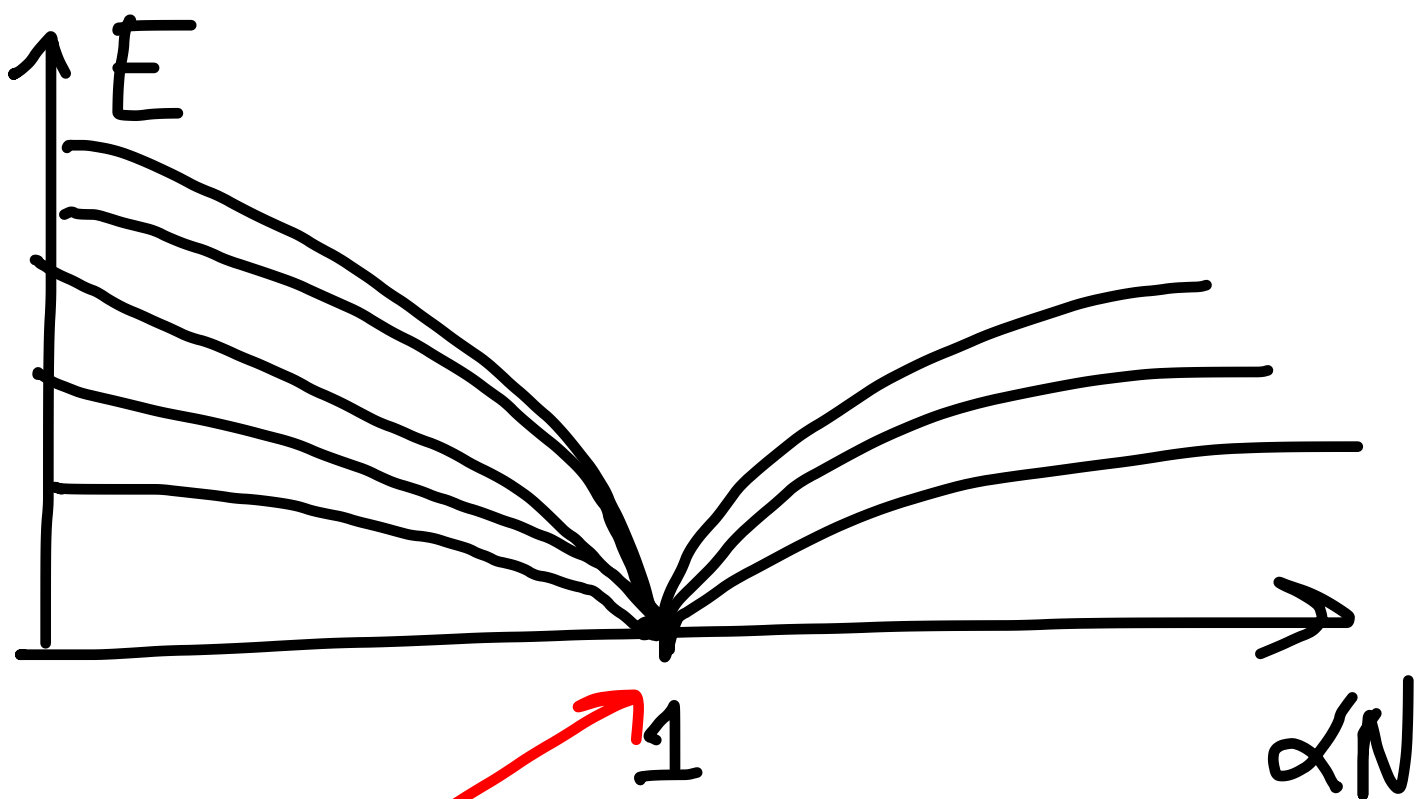
$$\alpha N < 1$$



$$\alpha N > 1$$

collapse

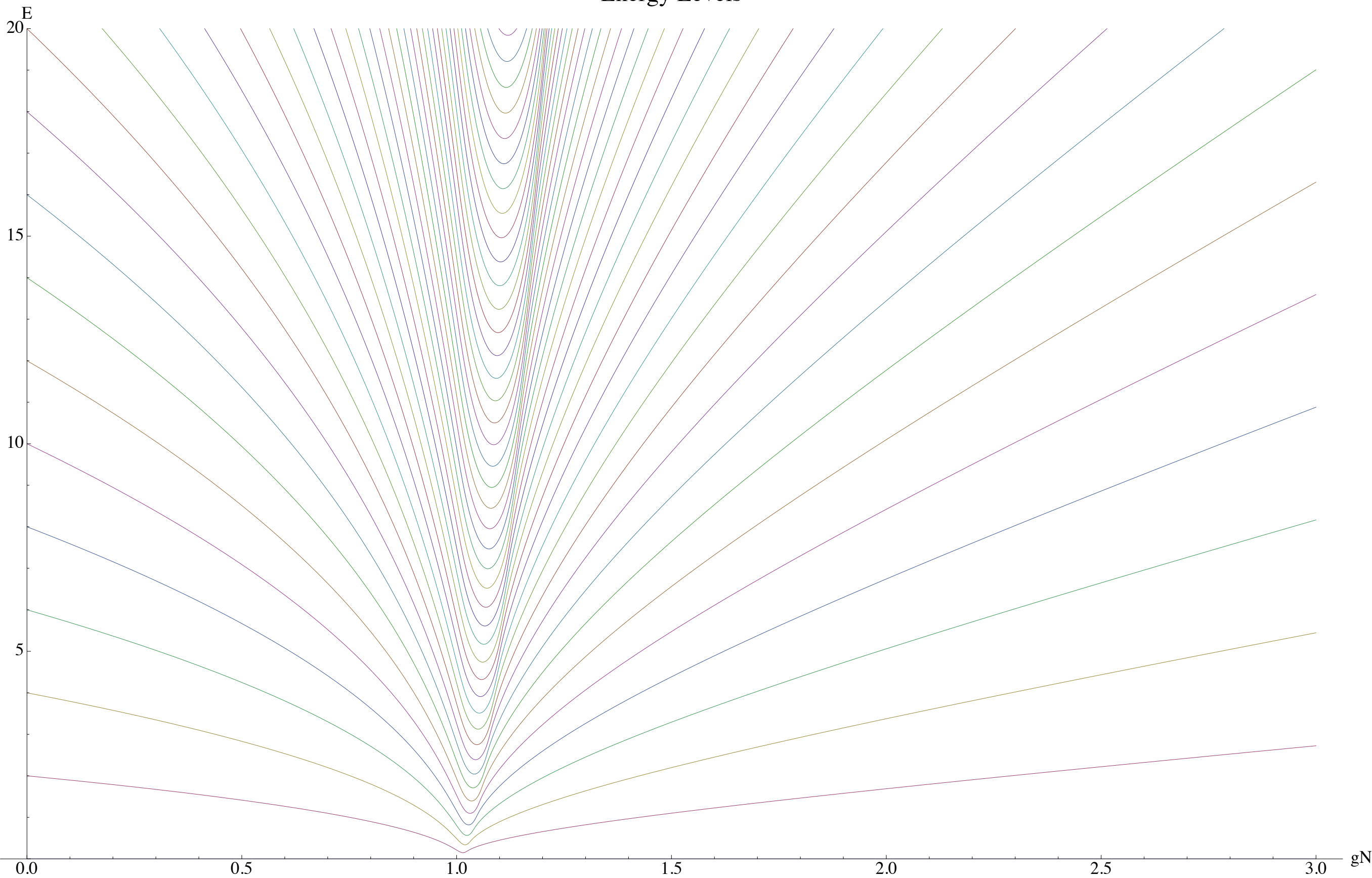
$\alpha N = 1$  ← Critical point of quantum phase transition



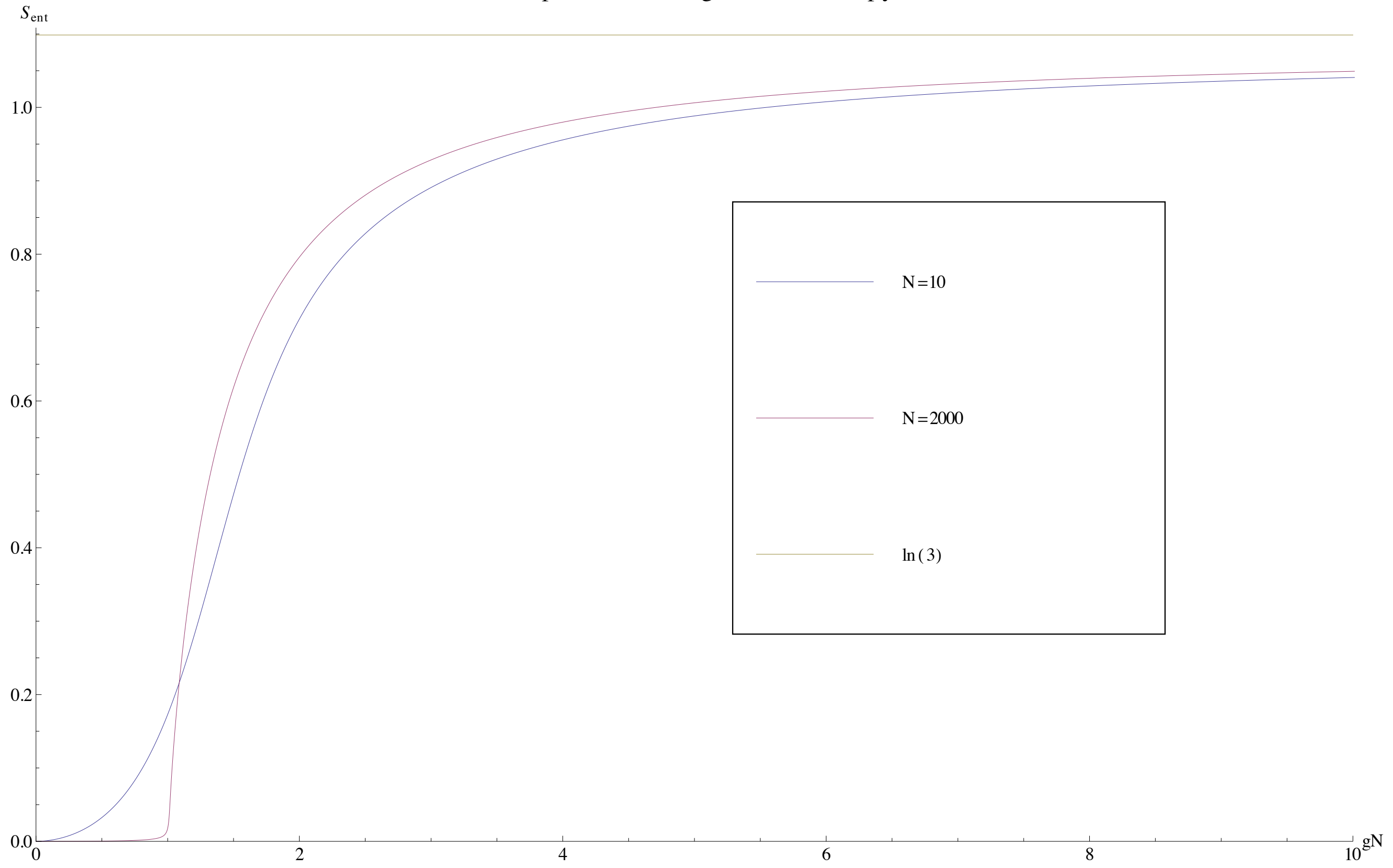
Many cheap (gapless) qubits!



Energy Levels



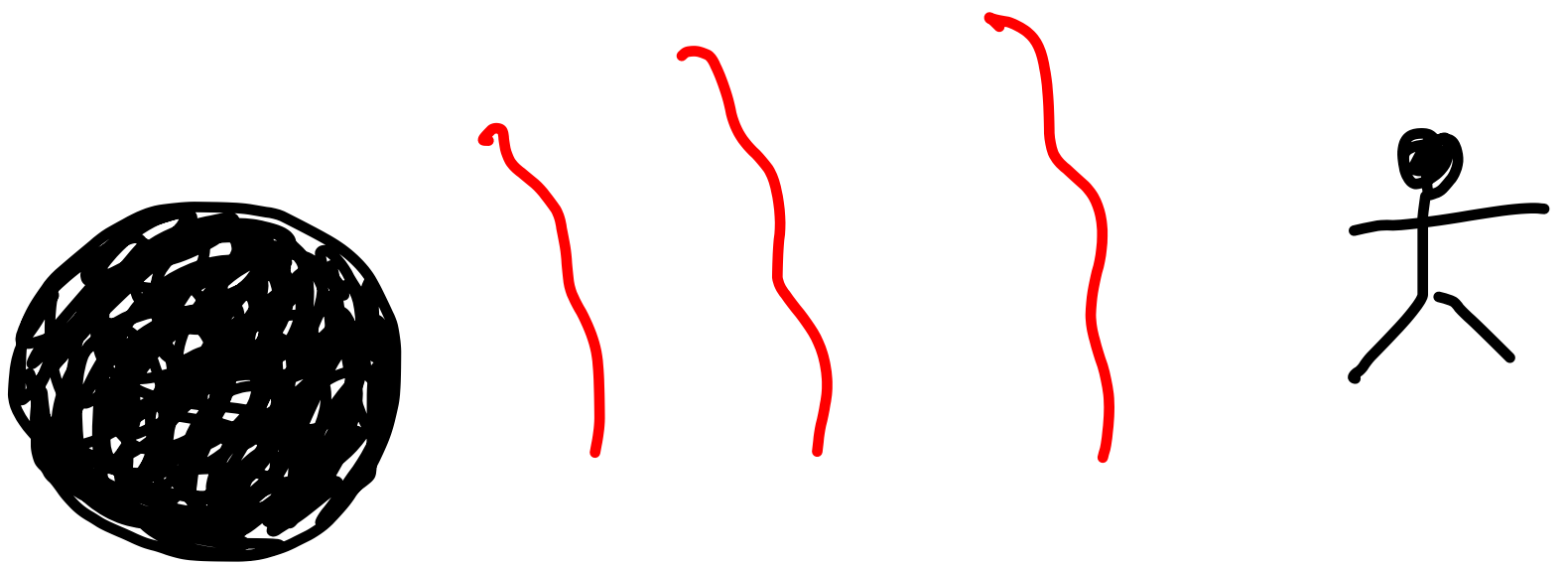
# One particle Entanglement Entropy



Black hole = Many  
gravitons at quantum  
critical point!

Self-sustained  
quantum criticality.

There is no paradox:  
 $t \rightarrow 0$  black hole indeed  
carries  $\infty$  information.

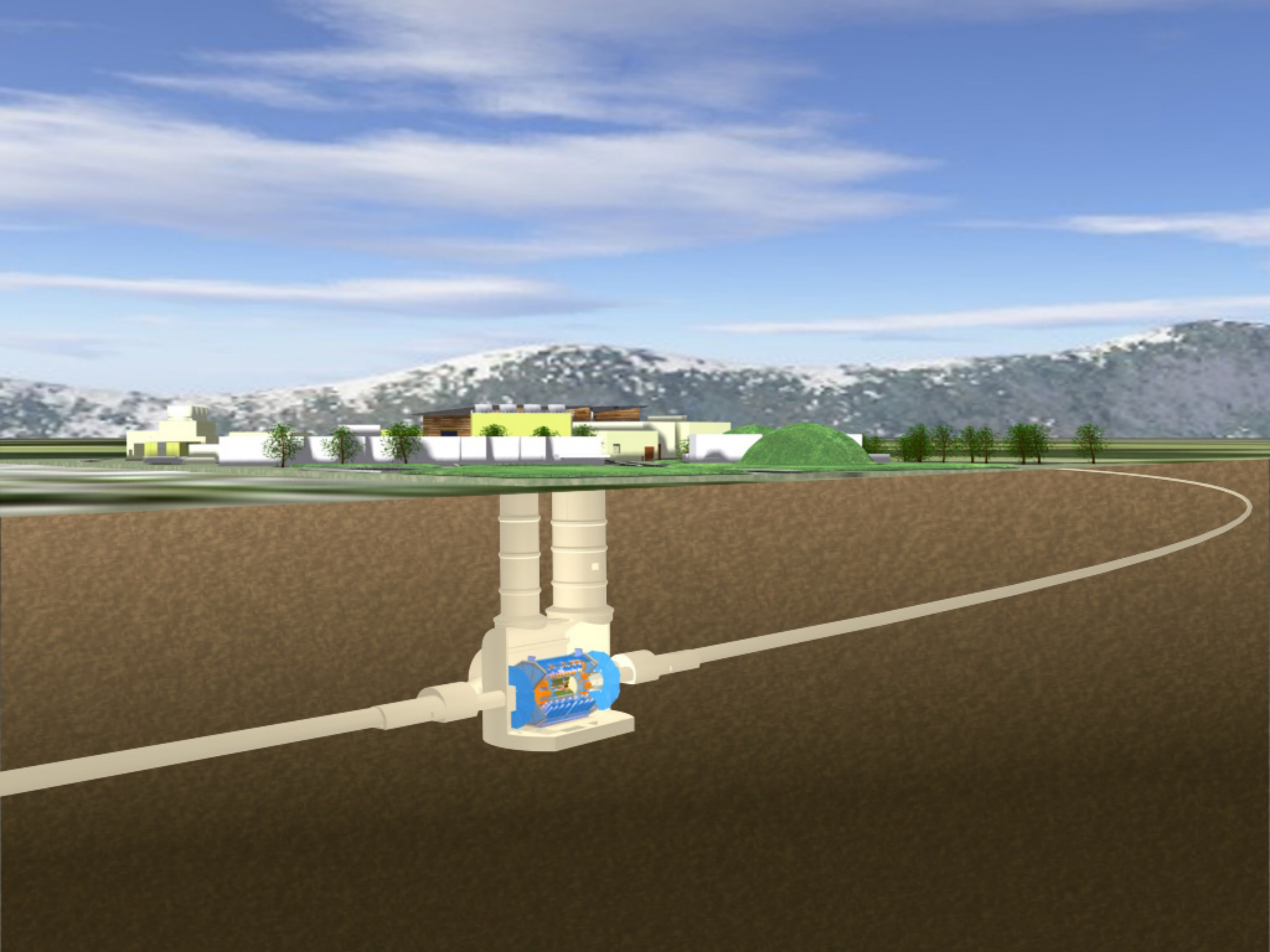


However it takes  
 $t = \infty$  to read it!

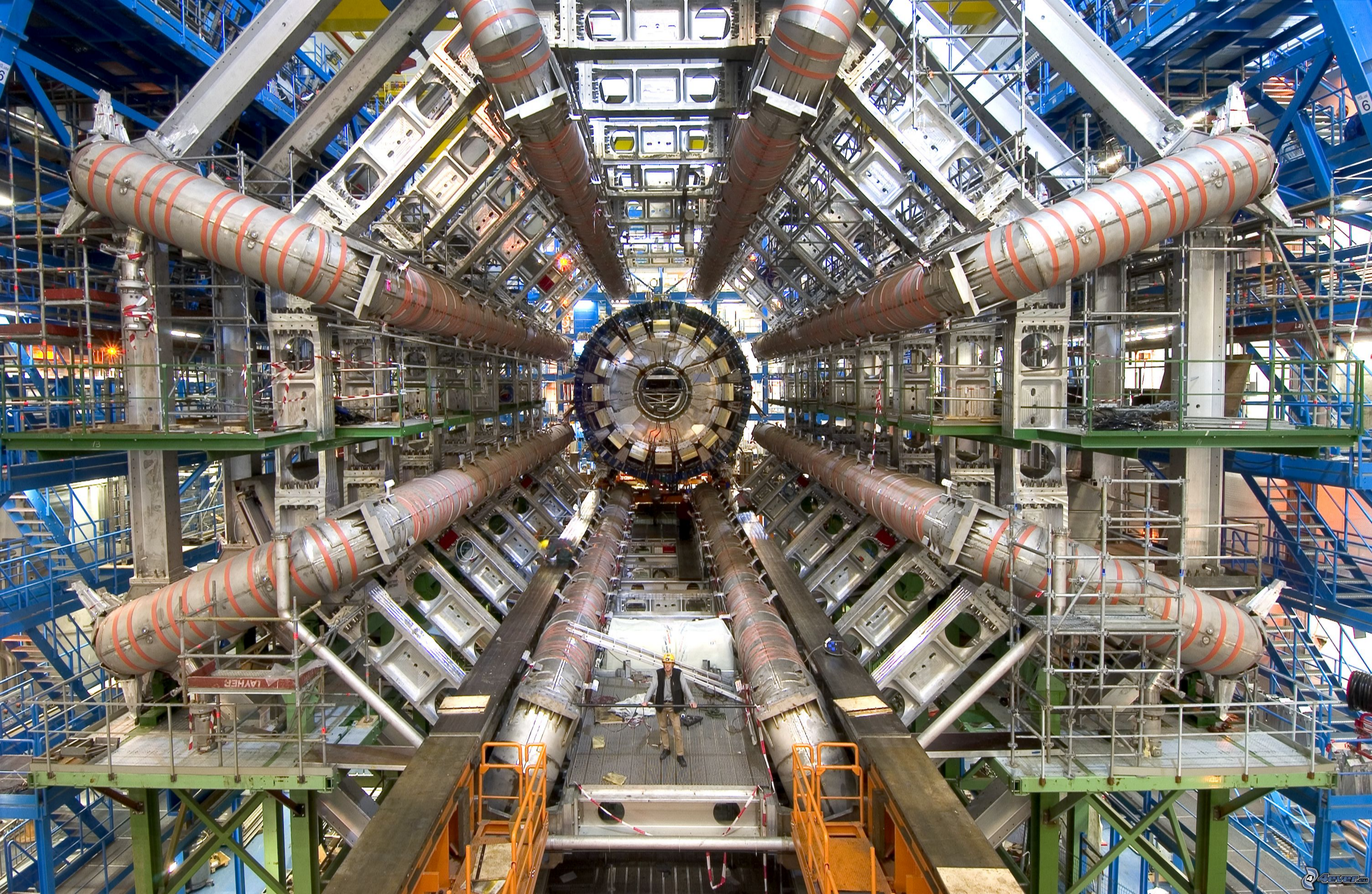
Experimental prospects:

⊛ Simulating black holes;

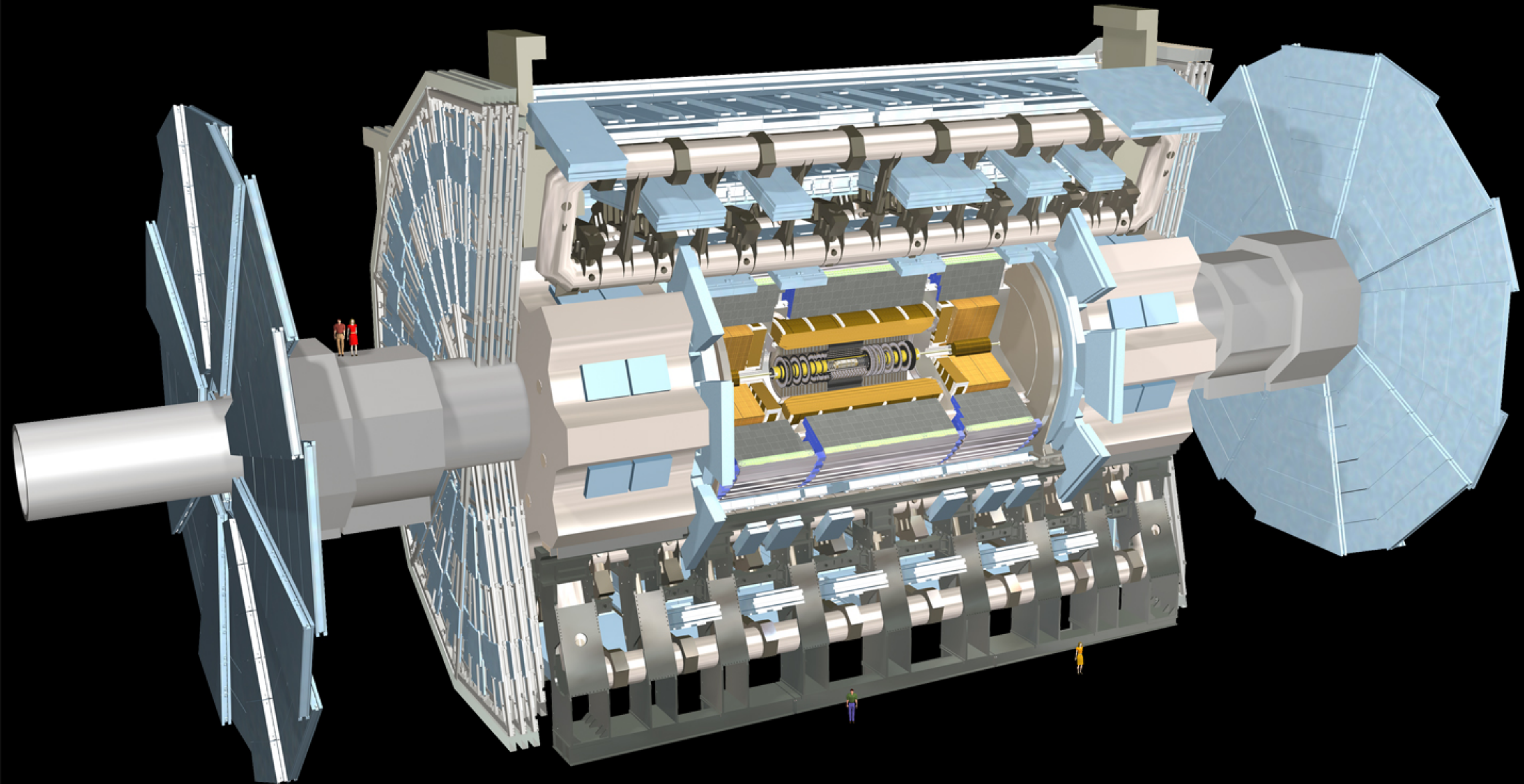
⊛ Black hole based  
quantum computing  
in labs!

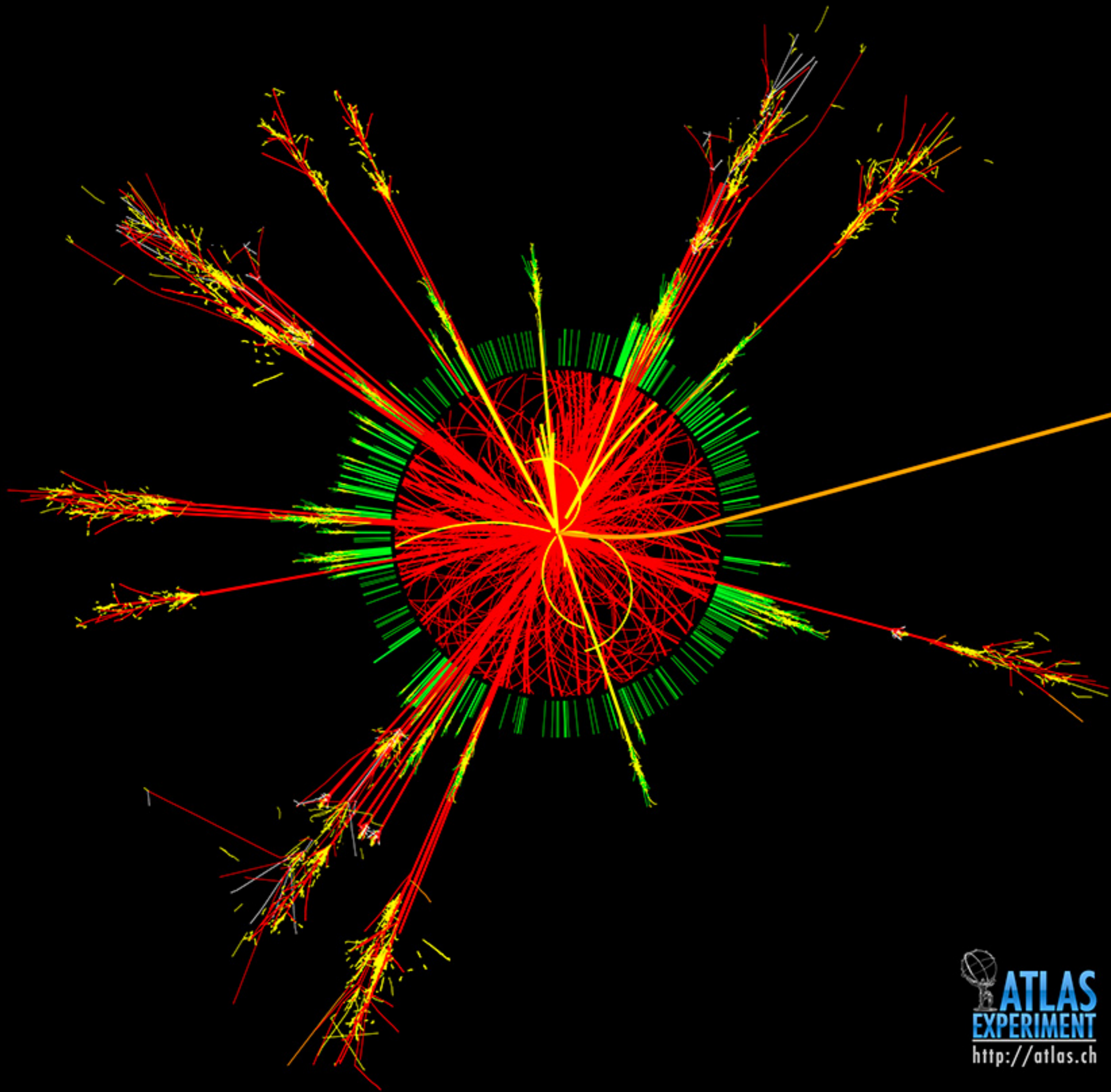












# Applications:

- ① \* Understanding Quantum criticality.
- ① \* Understanding black holes (in the sky and in Labs).
- ① \* Building black hole based quantum computers.
- ① \* Understanding UV-completion